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A Model for Medium Range Weather Forecasting

- Adiabatic Formulation -

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ABSTRACT

In this report the design of the finite difference scheme for ECMWF's first global model using the primitive equations in spherical co-ordinates is described. The spatial differencing scheme is formulated in accordance with methods which are now well established and maintains many of the more important integral constraints satisfied by the continuous forms of the equation. The maintenance of these constraints inhibits non-linear instability and ensures a more accurate simulation of exchange and cascade processes.

The results of an adiabatic integration to three days and an integration to ten days with a simple surface drag formulation and convective adjustment scheme are presented.

Although not discussed in this paper, a high resolution limited area version of the model has been designed. This model will be used primarily for experiments with the model's parameter station schemes for sub-grid scale phenomena.

1. Introduction

Meteorologists are faced with a wide, though unenviable, choice of schemes for the basic building blocks of a numerical weather prediction model, whether it is to be used for research or operational forecasting. spatial representation the main areas of choice are grid point schemes, pseudo-spectral or functional approaches or Galerkin techniques using either local (for example finite elements) or global (for example true spectral) For medium range weather forecasting not basis functions. only is the accuracy of the initial data important in order to provide an accurate description of the evolution of the atmosphere in the early stages of the forecast, but a proper representation, achieved mainly through sub-grid scale parameterisations, of all the important physical processes for the atmosphere is necessary to determine the development areas of new systems not present in the initial data. the area of physical parameterisation the variety and levels of sophistication are enormous. In this paper we confine ourselves to a description of a formulation for an adiabatic calculation. The formulation for the diabatic part will follow in another technical report.

For the first operational model and for most of our research for the next few years we have chosen a grid point representation with second order difference approximations for the spatial derivatives. The model can be used in global or hemispheric (with a solid wall boundary condition at the equator) domains. In addition a limited area version is available which uses a relaxation technique described by DAVIES (1976). limited area version the model's dependent variables are relaxed towards externally specified values within a narrow This technique provides possibly a better boundary zone. alternative to the solid wall equatorial boundary condition for hemispheric integrations. In this case the model could be relaxed within a narrow equatorial band towards a combination of persistence and climatology or even towards values obtained from a global forecast produced by a coarse mesh integration. A preliminary description of the limited area version is available in an internal report by KALLBERG (1977).

The spatial differencing scheme has been designed to preserve many of the important integral constraints satisfied by the continuous equations; conservation of mass, conservation of moisture, conservation of potential absolute enstrophy during potential vorticity advection by the horizontal wind field and conservation of total energy apart from sources, sinks and boundary fluxes. Maintenance of integral constraints is probably not strictly necessary for short range forecasting in which the concern is with the accuracy of the local description. However, for longer term integrations maintenance of these constraints is essential in order to prevent systematic errors and to maintain the In particular statistical structure of the atmosphere. conservation of potential enstrophy, or enstrophy in barotropic nondivergent flows, not only prevents non-linear instability,

ARAKAWA (1966) and SADOURNY (1975a), but it is an essential feature of 2-dimensional barotropic flows and controls the exchanges of energy between different scales. Failure to conserve enstrophy in 2-dimensional non-divergent flows leads to a spurious computational cascade of energy to small scales resulting in an energy catastrophe, SADOURNY (1975a). The use of ad hoc lateral diffusion schemes can control the false numerical cascade by directly removing energy at the small scale end of the spectrum, but a false energy cascade into these scales in combination with excessive lateral smoothing, which is usually necessary to stabilize schemes with this defect, enhances the total amount of energy dissipation and removes erroneously energy from the weather bearing systems. It is therefore necessary to describe the dynamics of the truncated system (BASDEVANT and SADOURNY (1975)) as correctly as possible and to accurately describe the transfers of energy (and enstrophy) across the boundary imposed by truncation. An essential ingredient is conservation of enstrophy during vorticity advection by the non-divergent part of the wind.

The vertical co-ordinate is pressure normalised by the surface pressure, a sigma system, which allows exact satisfaction of the kinematic boundary condition at the surface of the earth. The model can be used with any specified irregular vertical resolution. We have chosen to use a regular spatially staggered latitude/longitude grid in the horizontal. The time-stepping algorithm is based on the explicit leap-frog scheme together with a spatial filtering technique to overcome the severe stability restriction arising from the convergence of the meridians to the poles.

In section 2 we describe the continuous equations for the model and in sections 3 and 4 the spatial and temporal differencing schemes are described.

In the last section, section 5, we present results of adiabatic integrations and also of some integrations with a dry convective adjustment scheme and a simple surface stress formulation.

2. The governing equations in sigma co-ordinates

The simplest sigma (σ) co-ordinate system has been chosen for the first version of the model, namely the system proposed by PHILLIPS (1957), in which

$$\sigma = p/p_{s} \tag{1}$$

where p is pressure and $p_{\rm S}$ is the pressure at the earth's surface. The sigma co-ordinate vertical velocity is

$$\dot{\sigma} = d\sigma/dt$$
 (2).

At the top of the model atmosphere ($\sigma=0$) we have set $\dot{\sigma}=0$ which gives the upper boundary condition

$$(p_s \mathring{\sigma})_{\sigma=0} = 0.$$
 (3).

At the surface of the earth (σ =1) we have the kinematic boundary condition

$$(p_{s}\dot{\sigma})_{\sigma=1} = 0. \tag{4}.$$

Assuming the model atmosphere to be a perfect gas the hydrostatic equation can be expressed as

$$\frac{\partial \phi}{\partial \ln \sigma} = -RT \tag{5}$$

where R is the gas constant, which for the purpose of this paper takes the value for dry air. A list of the variables and constants used in this paper is given in Table 1. The mass continuity equation (note the density in this sigma system is proportional to $p_{\rm S}$) is

$$\frac{\partial p_{S}}{\partial t} + \frac{1}{a\cos(\theta)} \left\{ \frac{\partial}{\partial \lambda} (p_{S}u) + \frac{\partial}{\partial \theta} (p_{S}v\cos(\theta)) \right\} + \frac{\partial}{\partial \sigma} (p_{S}\dot{\sigma}) = 0$$
 (6).

This continuity equation is used to determine both på and $\frac{\partial p_S}{\partial t}$ Integrating equation (6) with respect to σ from $\sigma = 0$ to σ

and using the upper boundary condition (3) we obtain

$$\sigma \frac{\partial p_{s}}{\partial t} + p_{s} = - \begin{cases} \sigma \\ \frac{1}{a\cos(\theta)} \left\{ \frac{\partial}{\partial \lambda} (p_{s}u) + \frac{\partial}{\partial \theta} (p_{s}v\cos(\theta)) \right\} d\sigma \end{cases}$$
 (7).

On setting $\sigma = 1$ in (7) and using the kinematic boundary condition (4) we obtain the pressure tendency equation

$$\frac{\partial p_{S}}{\partial t} = -\int_{0}^{1} \frac{1}{a\cos(\theta)} \{ \frac{\partial}{\partial \lambda} (p_{S}u) + \frac{\partial}{\partial \theta} (p_{S}v\cos(\theta)) \} d\sigma$$
 (8)

 $p_S{}^{\dot\sigma}$ can now be determined from (7) and (8) by eliminating $\partial p_S/\partial t$. The zonal and meridional components of the momentum equations may be written

$$\frac{\partial u}{\partial t} - \frac{1}{\cos(\theta)} \operatorname{Zp}_{S} v \cos(\theta) + \frac{1}{\operatorname{acos}(\theta)} \frac{\partial}{\partial \lambda} (\phi + E) + \operatorname{RT} \frac{1}{\operatorname{acos}(\theta)} \frac{\partial}{\partial \lambda} (\ln p_{S}) + \frac{\partial}{\partial \sigma} = F_{u}$$
(9)

and

$$\frac{\partial v}{\partial t} + Zp_S u + \frac{1}{a} \frac{\partial}{\partial \theta} (\phi + E) + RT \frac{1}{a} \frac{\partial}{\partial \theta} (lnp_S) + \dot{\sigma} \frac{\partial v}{\partial \sigma} = F_V$$
 (10)

where Z is the potential absolute vorticity defined by

$$Z = \frac{1}{p_{S}} \left\{ f + \frac{1}{a\cos(\theta)} \left(\frac{\partial v}{\partial \lambda} - \frac{\partial}{\partial \theta} (u\cos(\theta)) \right) \right\}$$
 (11)

and E is the kinetic energy per unit mass given by

$$E = \frac{1}{2} \left(u^2 + \frac{1}{\cos(\theta)} v^2 \cos(\theta) \right)$$
 (12)

The forms of the non-adiabatic forcing on the right hand sides of equations (9) and (10), and other equations, are not given in this paper. The first law of thermodynamics can be written

$$\frac{\partial T}{\partial t} + \frac{1}{p_{s}} \left[\frac{1}{a\cos(\theta)} (p_{s}u \frac{\partial T}{\partial \lambda} + p_{s}v\cos(\theta) \frac{\partial T}{\partial \theta}) + p_{s}\dot{\sigma} \frac{\partial T}{\partial \sigma} - \frac{\kappa T\omega}{\sigma} \right] = Q$$
(13)

where ω is the pressure vertical velocity defined by

$$\omega = \frac{\mathrm{d}p}{\mathrm{d}t} = \frac{\mathrm{d}}{\mathrm{d}t} (p_{S\sigma}) = p_{S\dot{\sigma}} + \sigma(\frac{\partial p_S}{\partial t} + \frac{1}{\mathrm{acos}(\theta)}) (u\frac{\partial p_S}{\partial \lambda} + v\cos(\theta)\frac{\partial p_S}{\partial \theta}))$$

$$= p_S\dot{\sigma} + \frac{\sigma}{\partial t} + \frac{\sigma}{\mathrm{acos}(\theta)} \left\{ p_S u \frac{\partial}{\partial \lambda} (\ln p_S) + p_S v\cos(\theta)\frac{\partial}{\partial \theta} (\ln p_S) \right\}$$

$$(14)$$

The continuity equation for water vapour is

$$\frac{\partial q}{\partial t} + \frac{1}{p_{S}} \left[\frac{1}{a\cos(\theta)} (p_{S}u \frac{\partial q}{\partial \lambda} + p_{S}v\cos(\theta) \frac{\partial q}{\partial \theta}) + p_{S}o\frac{\partial q}{\partial \sigma} \right] = S.$$
 (15)

When $\dot{\sigma}$, ω and the non-adiabatic terms are determined, equations (6),(7),(8),(9),(10),(13) and (15) can be solved for the time changes of p_s , u, v, T and q and a forecast can be made.

p	pressure surface pressure
T	<pre>geopotential (g x height) temperature longitudinal/zonal velocity</pre>
V	latitudinal/meridional velocity
$Z = \frac{1}{p_{s}} \left\{ f + \frac{1}{a\cos(\theta)} \left(\frac{\partial v}{\partial \lambda} - \frac{\partial (u\cos(\theta))}{\partial \theta} \right) \right\}$	humidity mixing ratio potential absolute vorticity
$E = \frac{1}{2} \left(u^2 + \frac{1}{\cos(\theta)} v^2 \cos(\theta) \right)$	-kinetic energy/unit mass
$\dot{\sigma} = \frac{d\sigma}{dt}$	sigma vertical velocity
ac ac	pressure vertical velocity
σ= p/p _s	vertical co-ordinate
λ	latitude
$f = 2 \Omega \sin(\theta)$ F_u , F_r , Q,S	non-adiabatic effects - sources/sinks
	geopotential of the earth's surface
R	
$egin{array}{c} C_p & & & & & & & & & & & & & & & & & & &$	specific heat at constant pressure angular velocity of the earth acceleration due to gravity radius of the earth.

Table 1 : LIST OF VARIABLES AND CONSTANTS USED IN THE MODEL.

3. The spatial differencing scheme

The dependent variables are staggered in space but not in time; the distribution of variables is illustrated in figures 1a and 1b. The horizontal wind components, temperature and humidity mixing ratio are kept at the main levels, levels k, and the vertical velocity $\dot{\sigma}$ and geopotential ϕ are kept at levels $k+\frac{1}{2}$. The vertical spacing of the levels is arbitrary and may vary from one level to another, whereas we use a regular latitude/longitude co-ordinate system in the horizontal. horizontal grid is the ARAKAWA C grid and has been used by ARAKAWA, see ARAKAWA and LAMB (1976), for general circulation experiments and in the British Meteorological Office, BURRIDGE (1975), for operational short range forecasting. This grid has good dispersion properties for inertia gravity . waves, giving an accurate simulation of the geostrophic adjustment process. In addition the C grid avoids two gridlength computational noise in the terms governing the motion of pure gravity waves.

We define

$$\Delta \sigma_{\mathbf{k}} = \sigma_{\mathbf{k} + \frac{1}{2}} - \sigma_{\mathbf{k} - \frac{1}{2}}, \tag{16}$$

then
$$K$$

$$\sum_{k=1}^{\infty} \Delta \sigma_k = 1,$$
(17)

where K is the number of main levels (temperature levels). The upper and lower boundary conditions become

$$\left(p_{S}\dot{\sigma}\right)_{\frac{1}{2}} = 0 \tag{18}$$

and
$$(p_S \dot{\sigma})_{K+\frac{1}{2}} = 0$$
 (19)

The values of σ_k and $\sigma_{k+\frac{1}{2}}$ are not independently specified but are related through consistency relations arising from energy constraints imposed on the finite difference scheme. This point will be discussed towards the end of the section.

K-2 K-3/2 K-1 $K-\frac{1}{2}$ $K-\frac{1}{2}$ $K+\frac{1}{2}$ $K+\frac{1}{2$

Figure 1a: Vertical distribution of variables.

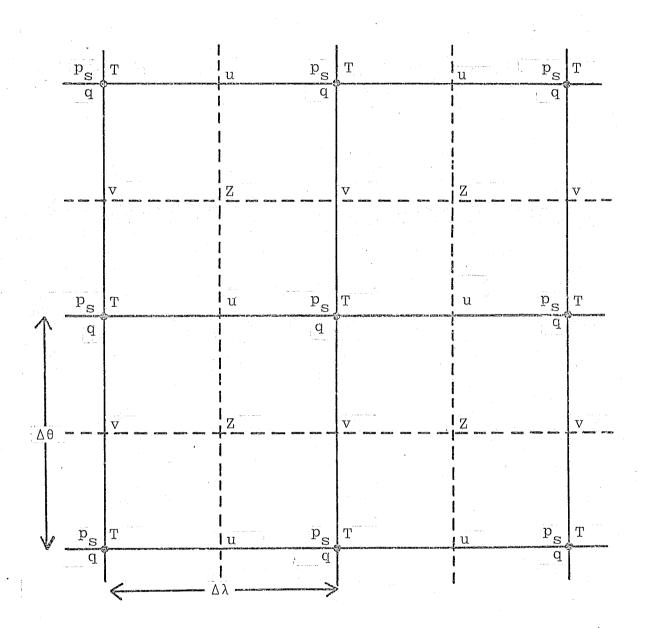


Figure 1b Horizontal distribution of variables. (The variables E, ϕ , ω and $\dot{\sigma}$ are kept at temperature points).

In order to set down the finite difference scheme adopted, the following well known notation has been used :-

$$\overline{A}^{X}(x) = \frac{1}{2} \left(A(x + \frac{\Delta x}{2}) + A(x - \frac{\Delta x}{2}) \right)$$

$$\overline{A}^{X}(x,y) = \overline{A}^{X}, y(x,y) = \frac{1}{4} \left(A(x + \frac{\Delta x}{2}, y + \frac{\Delta y}{2}) + A(x - \frac{\Delta x}{2}, y + \frac{\Delta y}{2}) + A(x - \frac{\Delta x}{2}, y - \frac{\Delta y}{2}) + A(x - \frac{\Delta x}{2}, y - \frac{\Delta y}{2}) \right)$$

$$\Delta_{X}A(x) = A(x + \frac{\Delta x}{2}) - A(x - \frac{\Delta x}{2})$$

$$\delta_{X}A(x) = \frac{\Delta_{X}A(x)}{\Delta_{X}}, \text{ for constant } \Delta x,$$

where Δx and Δy are grid lengths in the x and y directions respectively.

It is convenient to define 'mass flux' variables U and V by

$$\begin{array}{c}
U = \overline{p_S}^{\lambda} u \\
V = \overline{p_S}^{\theta} v
\end{array}$$
(20).

The hydrostatic equation is written

$$\frac{\Delta_{\sigma}\phi}{\Delta_{\sigma}\ell n\sigma} = -RT \tag{21}$$

and a vertical summation of this equation gives

$$\phi_{\mathbf{k}+\frac{1}{2}} = \phi_{\mathbf{S}} + \sum_{\ell=\mathbf{k}+1}^{\mathbf{K}} (\Delta_{\sigma} \ell n \sigma)_{\ell}$$
(22)

For the continuity equation and its vertically integrated forms, equations (6), (7) and (8), we have

$$\frac{\partial p_{s}}{\partial t} = \frac{1}{a\cos(\theta)} \left\{ \delta_{\lambda} U + \delta_{\theta} (V\cos(\theta)) \right\} + \frac{\Delta_{\sigma} (p_{s}\dot{\sigma})}{\Delta_{\sigma} \sigma} = 0$$
 (23)

$$\sigma_{\mathbf{k}+\frac{1}{2}} \frac{\partial \mathbf{p}_{\mathbf{S}}}{\partial \mathbf{t}} + (\mathbf{p}_{\mathbf{S}} \mathring{\sigma})_{\mathbf{k}+\frac{1}{2}} = -\frac{1}{\mathbf{a}\mathbf{cos}(\theta)} \sum_{\ell=1}^{K} \{\delta_{\lambda} \mathbf{U} + \delta_{\theta} (\mathbf{V}\mathbf{cos}(\theta))\}_{\ell} (\Delta_{\sigma} \sigma)_{\ell}$$
(24)

and

$$\frac{\partial p_{s}}{\partial t} = -\frac{1}{a\cos(\theta)} \sum_{\ell=1}^{K} \{\delta_{\ell} U + \delta_{\theta} (V\cos(\theta))\} (\Delta_{\sigma} \sigma)$$
(25)

The special forms of all the finite difference equations for points in the neighbourhood of, and at, the poles are given separately at the end of this section. The global integral (finite difference sum) of the right hand side of equation (25) vanishes (with appropriate polar boundary conditions) ensuring exact conservation of the model's total mass. The momentum equations are differenced as (leaving the non-adiabatic terms in symbolic form)

$$\frac{\partial \mathbf{u}}{\partial t} - \frac{1}{\cos(\theta)} \left[\text{ZV}\cos(\theta) \right] + \frac{1}{a\cos(\theta)} \delta_{\lambda} (\tilde{\phi}^{\mathsf{T}} \mathbf{E}) + \frac{R\overline{T}^{\lambda}}{a\cos(\theta)} \delta_{\lambda} (\ln p_{s}) + \frac{1}{p_{s}} \frac{\overline{p_{s}\dot{\sigma}^{\lambda}} \lambda_{G} \mathbf{u}}{\Delta_{\sigma}\sigma} = F_{u}$$
(26)

$$\frac{\partial v}{\partial t} + [ZU] + \frac{1}{a} \delta_{\theta} (\overline{\phi}^{\sigma} + E) + \frac{R\overline{T}}{a} \delta_{\theta} (\ln p_s) +$$

$$+\frac{1}{\overline{p_{s}}}\theta \frac{\overline{p_{s}\sigma} \Delta_{\sigma} v}{\Delta_{\sigma}\sigma} F_{v}$$
 (27)

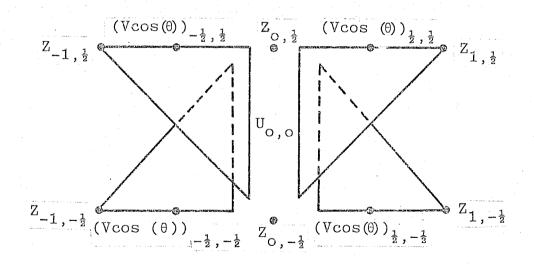
with

$$E = \frac{1}{2} (\overline{u^2}^{\lambda} + \frac{1}{\cos(\theta)} \overline{v^2 \cos(\theta)}^{\theta})$$
 (28)

and
$$Z = \frac{1}{p_{s}a\cos(\theta)} \lambda \theta \left\{ af\cos(\theta) + \delta_{\lambda} v - \delta_{\theta} \left(u\cos(\theta) \right) \right\}$$
 (29)

We have chosen to work with two finite difference schemes for the rotation terms $[\text{ZVcos}(\theta)]$ and [ZU]. The first conserves both energy and potential absolute enstrophy, $(p_8\text{Z}^2)$, and for non-divergent barotropic flow may be reduced to the energy/enstrophy conserving Jacobian designed by ARAKAWA(1966). The second conserves potential absolute enstrophy $(p_8\text{Z}^2)$ but not energy. The energy/enstrophy conserving scheme we have chosen cannot easily be expressed in terms of the operators introduced above; instead we use a geometric description (SADOURNY 1975b) in which the rotation terms are written:

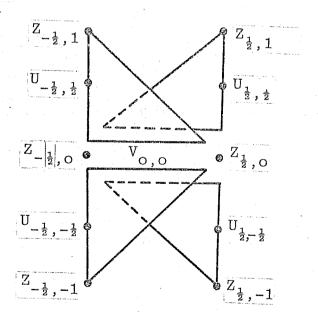
$$\begin{split} \left[\overline{Z} V \cos(\theta) \right]_{00} &= \frac{1}{4} \left[(V \cos(\theta))_{\frac{1}{2}, \frac{1}{2}} (Z_{0, \frac{1}{2}} + Z_{0, -\frac{1}{2}} + Z_{1, \frac{1}{2}}) / 3 \\ &+ (V \cos(\theta))_{-\frac{1}{2}, \frac{1}{2}} (Z_{0, \frac{1}{2}} + Z_{0, -\frac{1}{2}} + Z_{-1, \frac{1}{2}}) / 3 \\ &+ (V \cos(\theta))_{-\frac{1}{2}, -\frac{1}{2}} (Z_{0, \frac{1}{2}} + Z_{0, -\frac{1}{2}} + Z_{-1, -\frac{1}{2}}) / 3 \\ &+ (V \cos(\theta))_{\frac{1}{2}, -\frac{1}{2}} (Z_{0, \frac{1}{2}} + Z_{0, -\frac{1}{2}} + Z_{1, -\frac{1}{2}}) / 3 \end{bmatrix} \end{split}$$



and

$$\begin{bmatrix} ZU \end{bmatrix}_{0,0}^{=\frac{1}{4}} \begin{bmatrix} U_{\frac{1}{2},\frac{1}{2}} & (Z_{-\frac{1}{2},0}^{+}Z_{\frac{1}{2},0}^{+}Z_{\frac{1}{2},1}^{-}) & /3 \\ & +U_{-\frac{1}{2},\frac{1}{2}} & (Z_{-\frac{1}{2},0}^{+}Z_{\frac{1}{2},0}^{+}Z_{-\frac{1}{2},1}^{-}) & /3 \\ & +U_{-\frac{1}{2},\frac{1}{2}} & (Z_{-\frac{1}{2},0}^{+}Z_{\frac{1}{2},0}^{+}Z_{-\frac{1}{2},1}^{-}) & /3 \\ & +U_{-\frac{1}{2},-\frac{1}{2}} & (Z_{-\frac{1}{2},0}^{+}Z_{\frac{1}{2},0}^{+}Z_{-\frac{1}{2},-1}^{-}) & /3 \\ & +U_{\frac{1}{2},-\frac{1}{2}} & (Z_{-\frac{1}{2},0}^{+}Z_{\frac{1}{2},0}^{+}Z_{\frac{1}{2},-1}^{-}) & /3 \end{bmatrix}$$

$$(31)$$



For the enstrophy conserving scheme we have chosen

$$\left[\text{ ZVcos}(\theta) \right] = \overline{\text{Vcos}(\theta)}^{\text{Xy}} \overline{\text{Z}}^{\text{y}}$$
 (32)

and

$$\begin{bmatrix} zu \end{bmatrix} = \overline{v}^{xy} \overline{z}^{x} \tag{33}$$

The thermodynamic equation is used in the advective form

$$\frac{\partial T}{\partial t} + \frac{1}{p_{S}} \left\{ \frac{1}{a\cos(\theta)} \left(\overline{U_{\delta_{\lambda}}} \overline{T}^{\lambda} + \overline{V\cos(\theta)} \delta_{\theta} \overline{T}^{\theta} \right) + \frac{\overline{p_{S}\sigma \Delta_{\sigma}} T^{\sigma}}{\Delta_{\sigma}\sigma} - \left[\frac{\kappa T\omega}{\sigma} \right] \right\} = Q$$
(34)

with

$$\left[\frac{\kappa T_{\omega}}{\sigma}\right] = \frac{\kappa T}{\sigma} \left(\frac{\sigma \partial p_{S}}{\partial t} + p_{S} \sigma^{\sigma}\right) + \frac{\kappa}{a\cos(\theta)} \left(\frac{\overline{UT}^{\lambda} \delta_{\lambda}(\ln p_{S})}{(\ln p_{S})}\right) + \overline{V\cos(\theta)} \overline{T}^{\theta} \delta_{\theta} \\
\left(\frac{\pi T_{\omega}}{\partial t}\right) + \overline{V\cos(\theta)} \overline{T}^{\theta} \delta_{\theta} \\
\left(\frac{\pi T_{\omega}}{\partial t}\right)$$

Equation (34) may be rewritten in flux form using the finite difference continuity equation (23) and the finite difference rule

$$\Delta_{\mathbf{v}}(\mathbf{A}\overline{\mathbf{B}}^{\mathbf{X}}) = \overline{\mathbf{A}\Delta_{\mathbf{v}}\mathbf{B}}^{\mathbf{X}} + \mathbf{B}\Delta_{\mathbf{v}}\mathbf{A} \tag{36}.$$

The form chosen for the ω term, relation (35), ensures that transformations between kinetic energy and total potential energy cancel provided $1/\sigma$ is defined by

$$\frac{1}{\sigma} = \frac{\Delta_{\sigma} \ell n \sigma}{\Delta_{\sigma} \sigma} \tag{37}$$

Equation (37) is the consistency requirement relating values of σ at temperature levels (levels k) to those at $(p_S\sigma)$ levels (levels $k+\frac{1}{2}$). A relation such as (37) is not strictly necessary for energy conservation since energy conservation is guaranteed when κT in the first term on the right hand side of (35) is

replaced by $-\frac{\kappa}{R} \frac{\Delta_{\sigma} \phi}{\Delta_{\sigma} \sigma}$, a finite difference representation for $-\frac{\kappa}{R} \frac{\partial \phi}{\partial \sigma} (=\frac{\kappa T}{\sigma})$, and in this case we have

$$\left[\frac{\kappa T_{\omega}}{\sigma}\right] = -\frac{1}{C_{p}} \frac{\Delta_{\sigma} \phi}{\Delta_{\sigma} \sigma} \left[\frac{\sigma \partial p_{S}}{\partial t} + p_{S} \dot{\sigma}^{\sigma}\right] + \frac{\kappa}{a \cos(\theta)} - \frac{1}{U \overline{T}^{\lambda}} \delta_{\lambda} (\ln p_{S}) + V \cos(\theta) \overline{T}^{\theta} \delta_{\theta} (\ln p_{S})\right]$$
(38)

Relation (35) subject to (37) is of this form.

The moisture equation is

$$\frac{\partial \mathbf{q}}{\partial t} + \frac{1}{\mathbf{p}_{s}} \left\{ \frac{1}{\mathbf{a} \cos(\theta)} \left(\frac{1}{\mathbf{U} \delta_{\lambda} \mathbf{q}} + \frac{1}{\mathbf{V} \cos(\theta) \delta_{\theta} \mathbf{q}} \theta \right) + \frac{\mathbf{p}_{s} \dot{\sigma} \Delta_{\sigma} \mathbf{q}}{\Delta_{\sigma} \sigma} \right\} = \mathbf{S}$$
 (39)

which can be rewritten in a flux form and consequently moisture is conserved in the absence of sources and sinks.

3.1 Modification of the spatial differencing near and at the poles

We have chosen to keep temperatures at the poles as shown in figures 2a. and 2b. $P_s(=p_s,p)$ at the poles changes as a result of the meridional mass flux'V at all of the points on the latitude circle where the meridional velocity component v is carried. For the poles equation (23) is modified to give

$$\frac{\partial p_{s,p}}{\partial t} + (SIGN) \frac{a \Delta \lambda}{\varepsilon} \sum_{i=1}^{IX} (V\cos(\theta))_{p-\frac{1}{2},i} + \frac{\Delta_{\sigma}(p_{s}\dot{\sigma})}{\Delta_{\sigma}\sigma} p = 0$$
 (40)

where IX is the number of points on a latitude circle,

$$\varepsilon = (IX) \left(\frac{1}{2} a \Delta \lambda \cos\left(\frac{\pi}{2} - \frac{\Delta \theta}{2}\right)\right) \left(\frac{a \Delta \theta}{2}\right) \tag{41}$$

is the area represented by a polar grid point, and

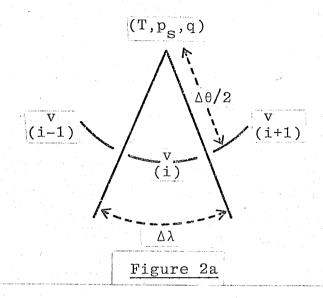
$$SIGN = -1$$
 for the North pole
 $SIGN = +1$ for the South pole (42).

The polar surface pressure tendency $\frac{\partial p_S}{\partial t}$, p and vertical velocity $(p_S^{\dot{o}})_p$ can be determined from vertical integrals (finite difference sums) of (40), leading to the two equations

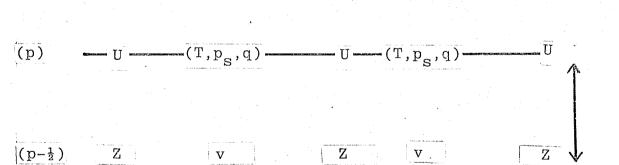
and
$$\frac{\sigma_{\mathbf{k}+\frac{1}{2}} \frac{\partial \mathbf{p}_{\mathbf{S}}, \mathbf{p}}{\partial t} + \left((\mathbf{p}_{\mathbf{S}} \dot{\sigma})_{\mathbf{p}_{\mathbf{k}+\frac{1}{2}}} \right) = -\sum_{k=1}^{K} \left((\mathrm{SIGN}) \frac{\mathbf{a} \Delta \lambda}{\varepsilon} \sum_{\mathbf{i}=1}^{N} (\mathrm{Vcos}(\theta))_{\mathbf{p}-\frac{1}{2}}, \mathbf{i} \right) (\hat{\sigma}^{\sigma}) \ell$$

$$\frac{\partial \mathbf{p}_{\mathbf{S}}}{\partial t} \mathbf{p} = -\sum_{k=1}^{K} \left((\mathrm{SIGN}) \frac{\mathbf{a} \Delta \lambda}{\varepsilon} \sum_{\mathbf{i}=1}^{IX} (\mathrm{Vcos}(\theta))_{\mathbf{p}-\frac{1}{2}}, \mathbf{i} \right) (\hat{\sigma}^{\sigma}) \ell$$

$$(44)$$



Distribution of variables near and at the North pole. (latitude / longitude grid).



 $(p+\frac{1}{2})$

 $(i-\frac{1}{2})$ (i) $(i+\frac{3}{2})$ (i+1) $(i+\frac{3}{2})$

Figure 2b

Distribution of variables near and at the North pole.



mirror values

North pole.

 $\Delta\theta/2$

The finite difference equation(27), the v equation, may be used to calculate $V_{p-\frac{1}{2},i}$ provided 'mirror values' $Z_{p+\frac{1}{2},i+\frac{1}{2}}$, the polar zonal mass flux, $U_{p,i}$ and kinetic energy /unit mass E_p are defined. We have chosen

$$Z_{p+\frac{1}{2},i+\frac{1}{2}} = Z_{p-\frac{1}{2},i+\frac{1}{2}}$$
 (45)

$$E_{p} = \frac{1}{1 \times 1} \frac{1}{\frac{1}{2} \cos \left(\frac{\pi}{2} - \frac{\Delta \theta}{2}\right)} \sum_{i=1}^{1 \times 1} \frac{1}{2} \left(v_{p-\frac{1}{2},i}\right)^{2} \cos \left(\frac{\pi}{2} - \frac{\nabla \theta}{2}\right)$$

$$(46)$$

(NB. $\frac{1}{2} \cos \left(\frac{\pi}{2} - \frac{\Delta \theta}{2} \right)$ is the 'cos(θ)' for the polar cap).

The zonal mass flux $U_{p,i}$ at the poles is not a true prognostic variable and is determined from finite difference representations of the polar boundary conditions

e and is determined from finite difference representations polar boundary conditions
$$\frac{\partial U}{\partial \lambda} + \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\lambda d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}{\partial \theta} \left(V \cos(\theta) \right) d\theta = \frac{1}{2\pi} \int_{0}^{2\pi} \frac{\partial}$$

which take the forms

$$\begin{cases}
U_{p,i+\frac{1}{2}} - U_{p,i-\frac{1}{2}} \frac{a\Delta\theta}{2} + (SIGN)V_{p-\frac{1}{2},i} \cos(\theta_{p-\frac{1}{2}}) a\Delta\lambda = \\
= (SIGN) \frac{1}{IX} \sum_{i=1}^{IX} V_{p-\frac{1}{2},i} \cos(\theta_{p-\frac{1}{2}}) a\Delta\lambda
\end{cases} (48)$$

and

$$\sum_{i=1}^{IX} U_{p,i+\frac{1}{2}} = 0$$

respectively.

The choice of the mirror values of Z 'beyond' the poles (relation (45)) ensures conservation of potential absolute vorticity and potential absolute enstrophy by both forms chosen to represent the rotation terms, but energy conservation for the energy/enstrophy conserving scheme, (30) and (31), is limited by boundary effects at the poles resulting in boundary contributions from the rotation terms. These boundary terms are in fact very small and vanish for $Z_{p-\frac{1}{2},i+\frac{1}{2}}$ independent of i (Z constant on the

If we let the poles be v points (Z points, with Z = constant on the boundaries) with v at the poles restricted to a zonal wave number one component and allow cross-polar differencing where necessary, the energy/enstrophy conserving scheme, (30) and (31), can be designed to be strictly energy and potential absolute enstrophy conserving.

The equations for T_p and q_p are

$$\frac{\partial T_{p}}{\partial t} + \frac{1}{p_{s}} \left\{ (SIGN) \frac{a\Delta\lambda}{\varepsilon} \cos(\theta_{p-\frac{1}{2}}) \sum_{i=1}^{IX} (V_{p-\frac{1}{2},i} \overline{T}_{p-\frac{1}{2},i}^{\theta}) + \frac{P_{s\sigma}^{\circ} p_{s\sigma} T_{p}}{\Delta_{\sigma}^{\sigma}} - \left[\frac{\kappa T_{\omega}}{\sigma} \right] \right\} = Q_{p}$$
(49)

and

$$\frac{\partial q_{p}}{\partial t} + \frac{1}{p_{s}} \left\{ (SIGN) \frac{a\Delta\lambda}{\varepsilon} \cos(\theta_{p-\frac{1}{2}}) \sum_{i=1}^{IX} (V_{p-\frac{1}{2},i} \overline{q}_{p-\frac{1}{2},i}^{\theta}) + \frac{(\overline{p_{s\sigma}})_{p\Delta\sigma}q_{p}}{\Delta_{\sigma}\sigma} \right\} = S_{p}$$
(50)

with

$$\left[\begin{array}{c} \frac{\kappa T_{\omega}}{\sigma} \end{array}\right]_{p} = -\frac{1}{C_{p}} \frac{\Delta_{\sigma} \phi_{p}}{\Delta_{\sigma} \sigma} \left(\sigma \frac{\partial \overline{p_{S}} + p_{S}}{\partial t} + p_{S}\right)^{\sigma} + \frac{\kappa}{a^{\frac{1}{2}} \cos(\theta_{p-\frac{1}{2}})} \frac{1}{IX} \times$$

with or without the constraint (37).

The basically second order centred spatial differencing scheme described in this section endows the model with the following properties, subject only to time truncation, non-adiabatic effects and boundary terms:-

- (a) The total mass of the model is conserved.
- (b) Kinetic energy (for scheme 31 and 32) p_ST^2 , p_Sq^2 , moisture q and total potential energy are individually conserved under advective processes.
- (c) Potential absolute enstrophy, p_8Z^2 , and potential vorticity are conserved by the horizontal advection terms.
- (d) The total energy is conserved with the correct conversion between kinetic and total potential energy.

4. The time-stepping scheme

Let \underline{x} be a vector whose components are all the grid point values of the model's dependent variables, then we may formally write

$$\frac{\partial x}{\partial t} + \underline{\underline{A}}x + \underline{\underline{S}}x = 0 \tag{52}$$

where $\underline{\underline{A}}$ and $\underline{\underline{S}}$ are non-linear finite difference operators representing the adiabatic terms in the governing equations (the terms set out in the last section) and the non-adiabatic terms respectively. Our present algorithm is based on a leapfrog scheme for the dynamics and a forward step for the remaining terms (some processes such as vertical diffusion may have to be included using implicit techniques to prevent numerical instability). The integration scheme is

$$\underline{x} (t+\Delta t) = \underline{x} (t-\Delta t) - 2\Delta t \underline{A}\underline{x} (t) - 2\Delta t \underline{S}\underline{x} (t-\Delta t)$$
 (53)

where the overbar denotes a linear time-filtering operator such that

$$\frac{\overline{x}}{x}(t-\Delta t) = \underline{x}(t-\Delta t) + \alpha(\overline{x}(t-2\Delta t) - 2x(t-\Delta t) + x(t))$$
(54)

a typical value of α being 0.005. The properties of this filter are well known, ASSELIN (1972), and its purpose is to inhibit the growth of the spurious computational mode associated with the leap-frog scheme.

In order to overcome the severe stability restrictions arising from the convergence of the meridians to the poles, a latitudinally dependent spatial filtering operator, see ARAKAWA and LAMB (1976), is applied either selectively or to all the terms in the finite difference equations. Our filter can be formally expressed as the finite convolution operator

$$\underline{\underline{F}} = \underline{\underline{\Phi}}^{-1} \underline{\Lambda} \underline{\underline{\Phi}} \tag{55}$$

where $\underline{\Phi}$ is a finite Fourier transform along a line of latitude, Φ^{-1} its inverse and

$$\underline{\underline{\Lambda}} = \underline{\underline{\Lambda}}(\theta) = \operatorname{diag}(\underline{\Lambda}_{1}(\theta), ----, \underline{\Lambda}_{N}(\theta)),$$

where Λ_n (6) is the reduction factor for the n^{th} Fourier mode of the term being filtered. For linear systems this filter when applied to all terms (total tendency filtering) reduces the phase speeds of wave solutions of the equations. The filtering scheme has been designed so that the stability criterion for the finite difference scheme can be based on the east-west grid length at some fixed latitude, θ^O N, and the filter is then applied polewards of θ^O N and θ^O S in the Northern and Southern hemispheres respectively. Experiments with barotropic models using a number of filtering techniques will be described in another technical report.

5. Results

The integrations described in this section are based on initial data for 00 GMT 1st March 1965. The initial data set was constructed by interpolation from data on the GFDL N24 modified Kurihara grid, MIYAKODA (1974). Our integrations which included topography were made with the global energy/enstrophy scheme, (30) and (31) with the following resolution:-

$$\Delta\lambda \left(\frac{180^{\circ}}{\pi}\right) = \Delta\theta \left(\frac{180^{\circ}}{\pi}\right) = \left(\frac{90^{\circ}}{24}\right) = 3.75^{\circ} (\text{approximately N24})$$

nine temperature levels (the GFDL levels) with

$$\sigma_{k} = s_{k}^{2} (3-2s_{k})$$
 $k = 1, ---, 9$
 $s_{k} = (2_{k-1}) / 18$

and

$$\sigma_{k+\frac{1}{2}} = s_{k+\frac{1}{2}}^{2} (3-2s_{k+\frac{1}{2}})$$

$$s_{k+\frac{1}{2}} = 2k/18$$

and with the relation (38) for $\left[\frac{\kappa T\omega}{\sigma}\right]$.

The time-step was 5 minutes and no time smoothing (α = 0) or lateral diffusion was used. The spatial filter (53) was used for $/\Theta$ $(\frac{180^{\circ}}{\pi})/>450$ with the reduction factors, ARAKAWA and

LAMB (1976), defined as

$$\begin{array}{l} \Lambda_{\rm n}(\theta) = \frac{\cos(\theta^{\rm o})}{\cos(45^{\rm o})} & \frac{1}{\sin\left(\frac{{\rm n}\Delta\lambda}{2}\right)} & {\rm if} \left(\frac{\cos(\theta^{\rm o})}{\cos(45^{\rm o})}, \frac{1}{\sin\frac{{\rm n}\Delta\lambda}{2}}\right) \\ = 1 & {\rm if} \left(\frac{\cos(\theta^{\rm o})}{\cos(45^{\rm o})\sin\frac{{\rm n}\Delta\lambda}{2}}\right) \end{array} \geqslant 0$$

The model's initial data is illustrated in figures (3) and (4) by 500 mb and 1000 mb. height charts; NMC analyses for this time are shown in figures (5) and (6). An adiabatic forecast to 72 hours together with NMC analyses are illustrated (500 mb and 1000 mb height charts) in figures (7) to (18). It is not our intention to present a detailed analysis of the model's forecasting skill in this paper; its performance with the complete GFDL physical parameterisation scheme is demonstrated in a forthcoming technical report by GAUNTLETT et al (1977). The general flow pattern of the 500 mb forecasts compares well

with the analyses though the intensities of the individual features are quite different. At 1000 mb the systems are much too intense, typical of integrations with no surface friction. The wind field at the lowest levels in the model grew rapidly (within the first 12 hours) to about twice the original values accompanied by the intensification of surface features. The energy loss in this adiabatic experiment was less than 5×10^{-7} % per time-step. By 72 hours many grid points, particularly over topography, have become seriously statically unstable and the integration became unstable before four days.

A 10 day forecast which included a simple frictional drag scheme at the surface of the earth and a dry convective adjustment scheme is illustrated in figures (19) to (24). The surface stress for this integration was

$$\underline{\tau} = C_D \left(|\overline{V}^{\lambda}| u, |\overline{V}^{\theta}| v \right)$$

with
$$|V| = \sqrt{u^{2}} + \frac{1}{\cos(\theta)} \sqrt{v^{2} \cos \theta} = 0$$

and $C_{\rm D}$ (the non-dimensional drag coefficient) = 2 x 10⁻³. The major difference between the 72 hour forecasts, figures (11),(17),(20) and (23) is the large difference in the magnitude of the low level winds and the intensity of the surface features. The 10 day forecast has been included to demonstrate the stability of the model in the absence of ad hoc lateral smoothing schemes, though it is obvious from the 10 day 1000 mb chart, figure (24), there is a significant amount of noise at the limits of resolution of the model.

As mentioned above the next stage in the development of this model is to combine the adiabatic scheme of sections 3 and 4 with the GFDL physical parameterisation package and to evaluate its forecasting potential compared with that of other models available at ECMWF.

Acknowledgements

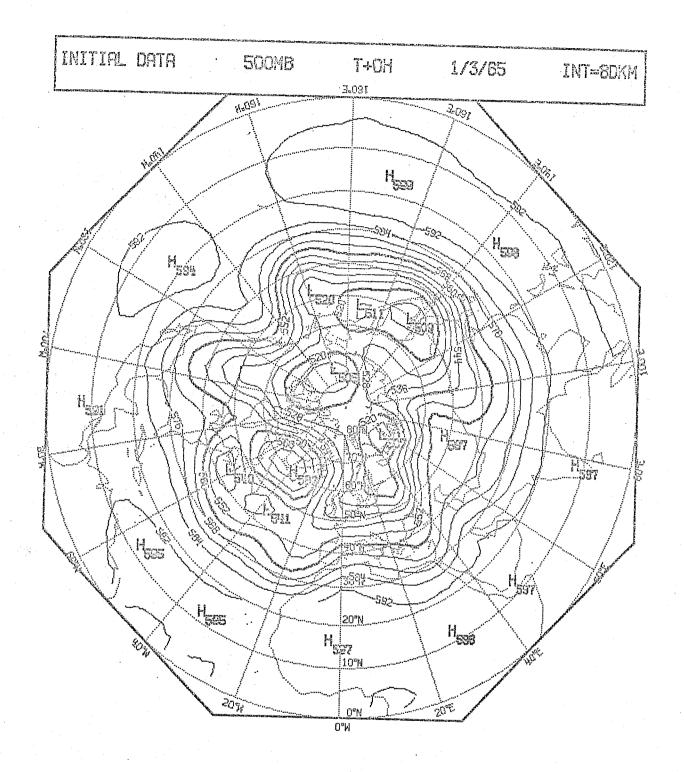
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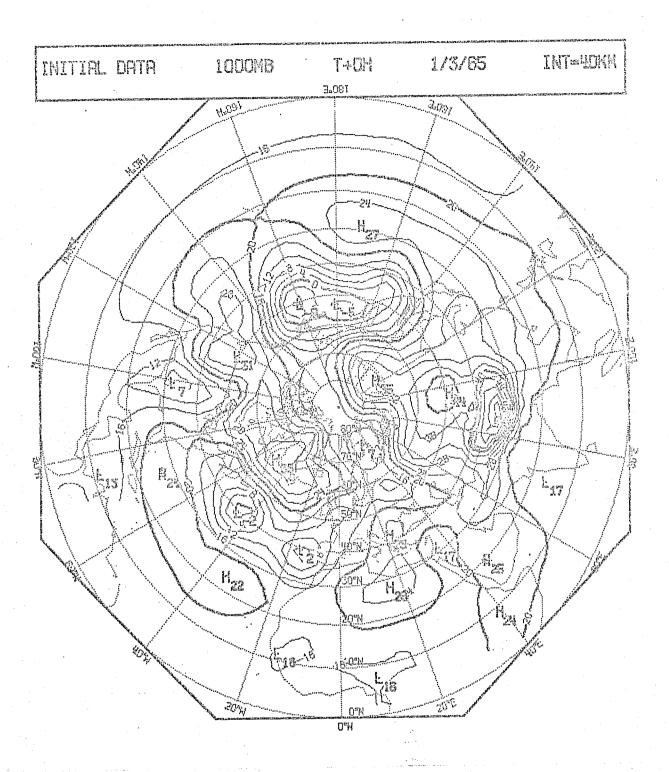
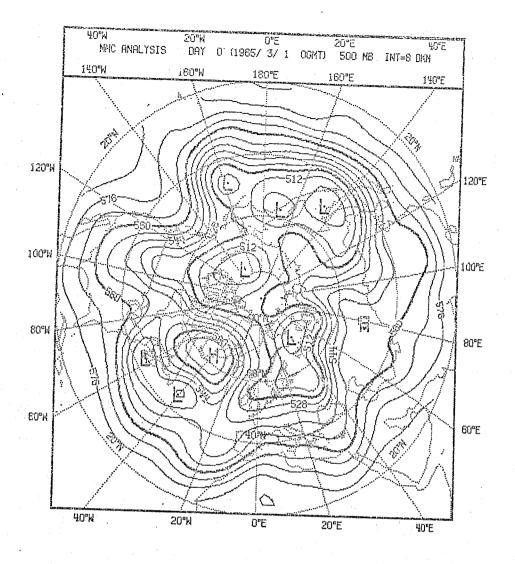
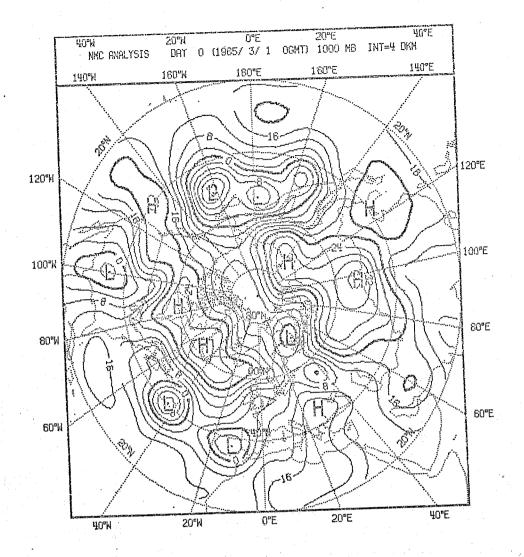
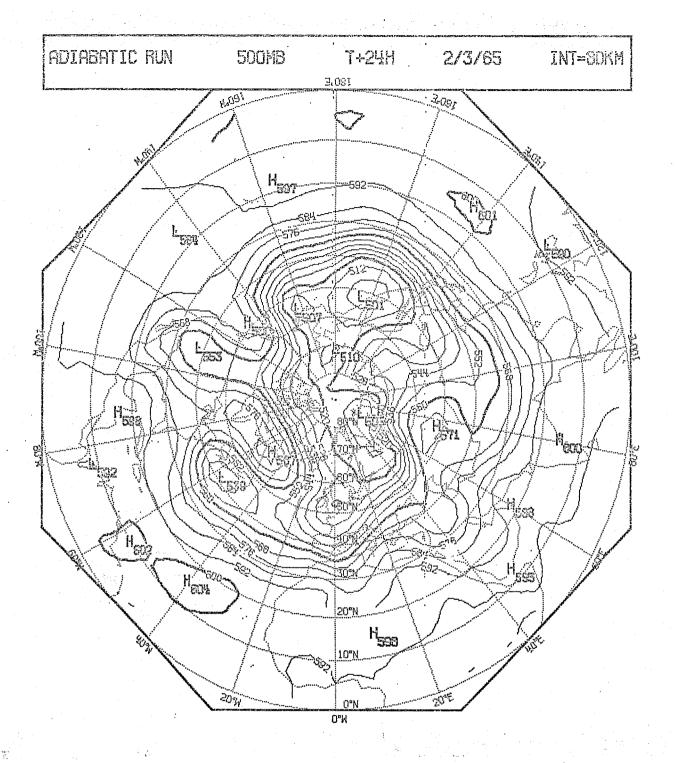
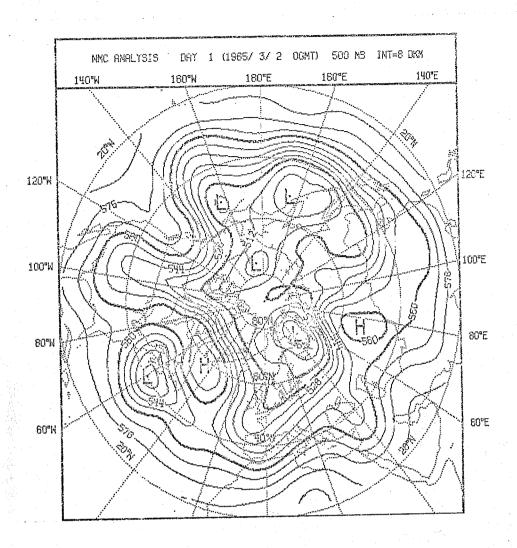


FIGURE 4









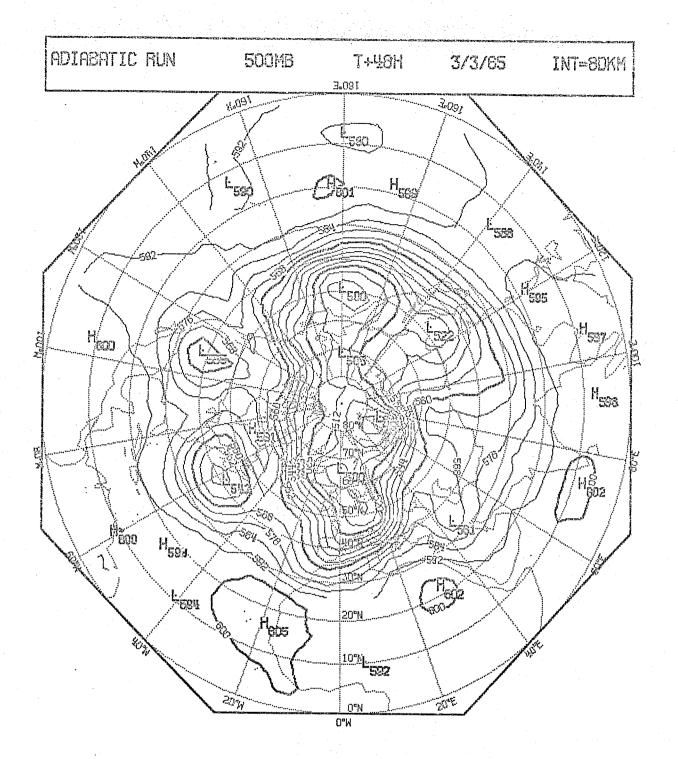
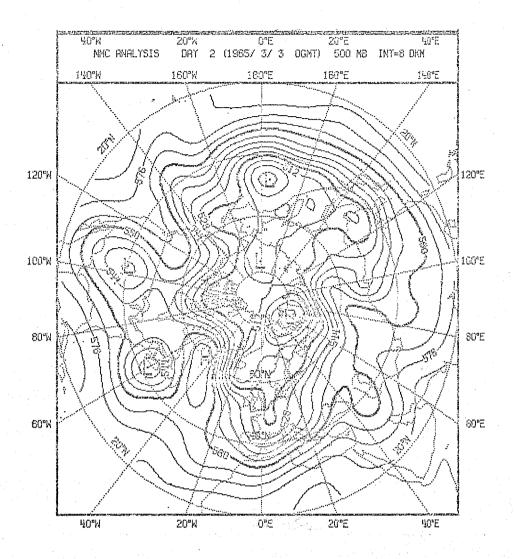
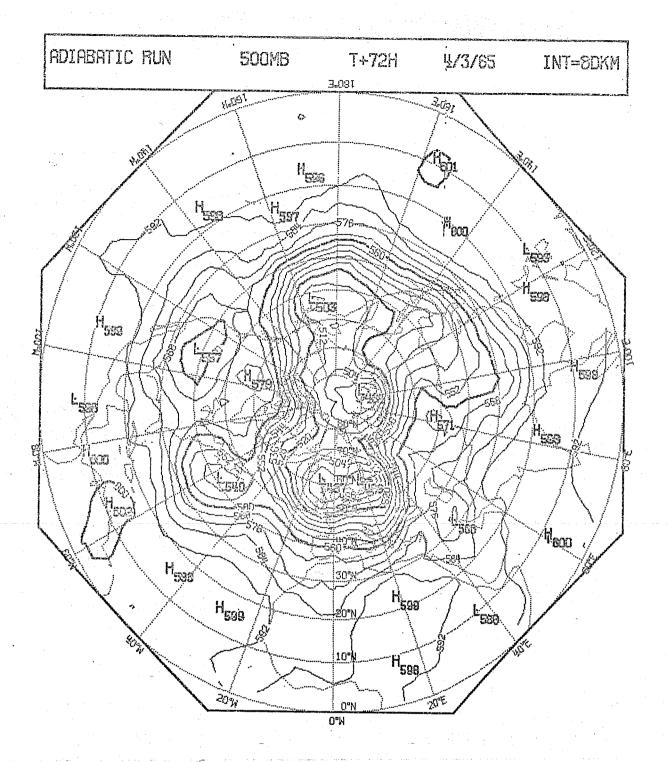
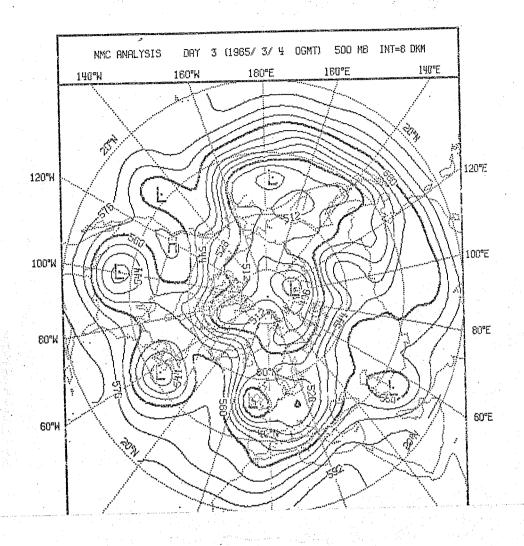




FIGURE 9







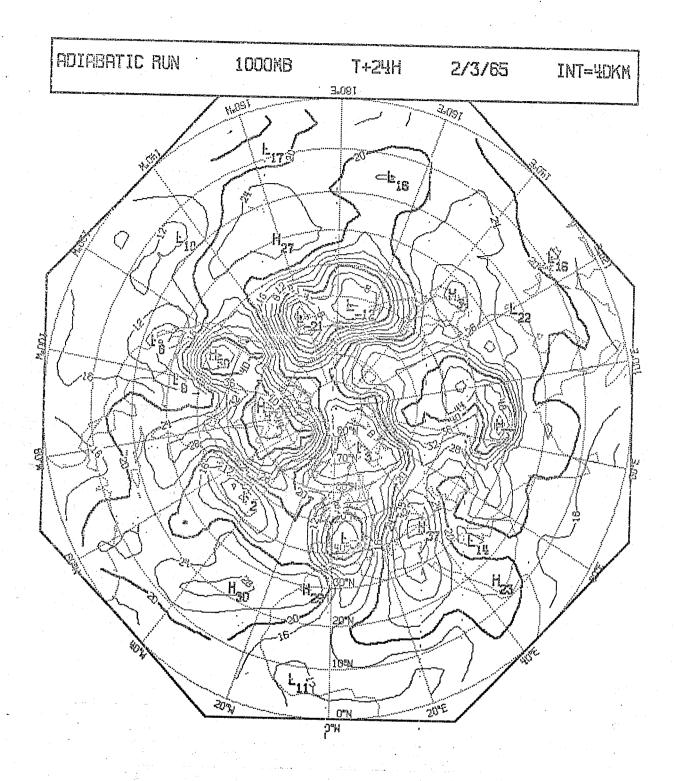
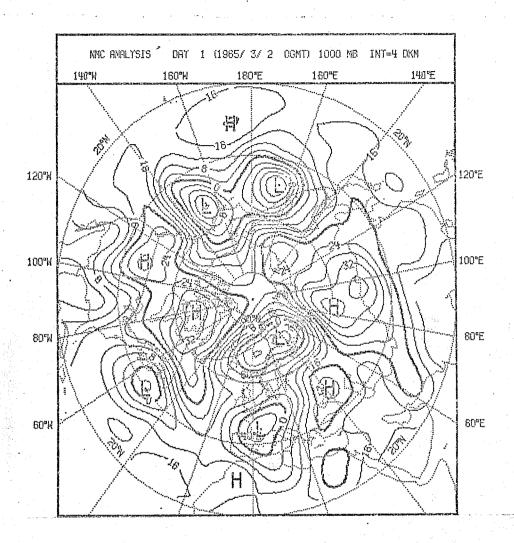
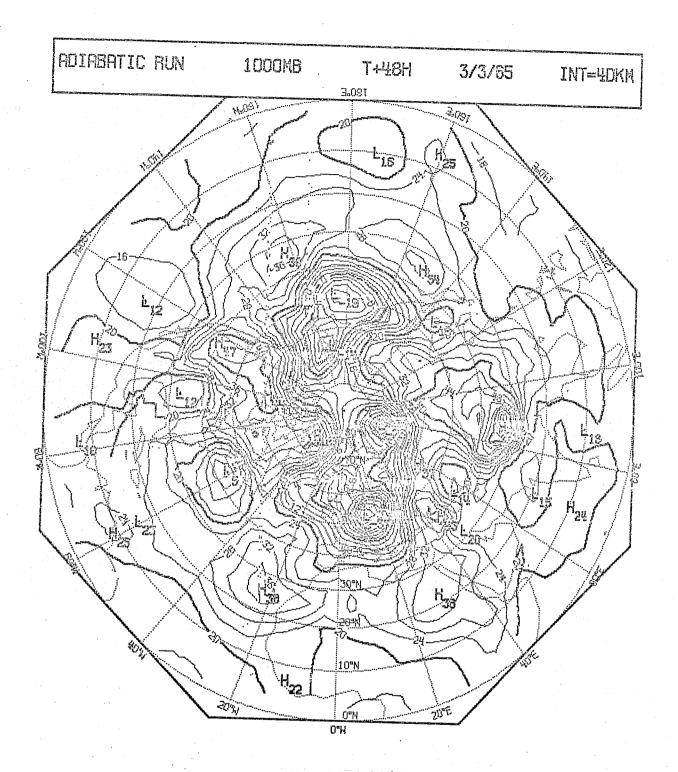
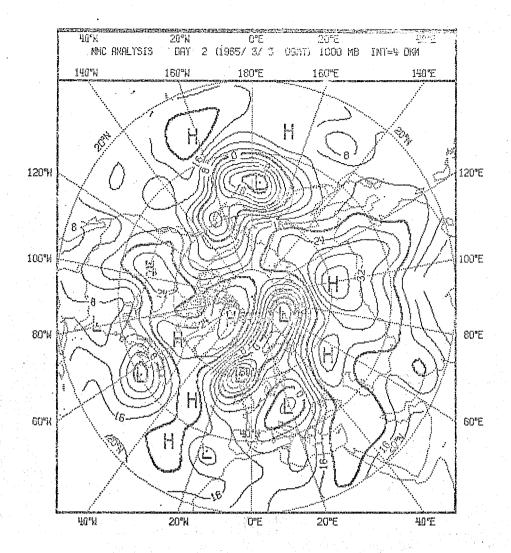
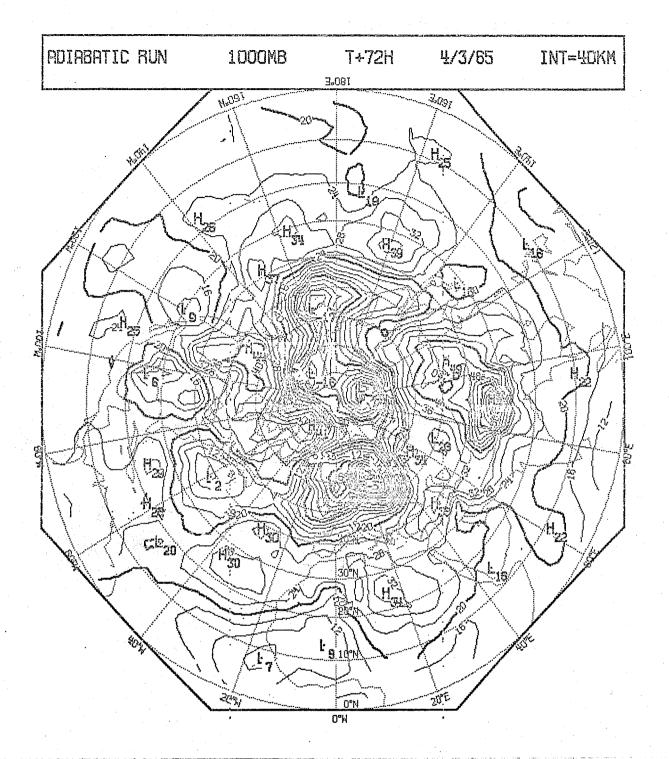


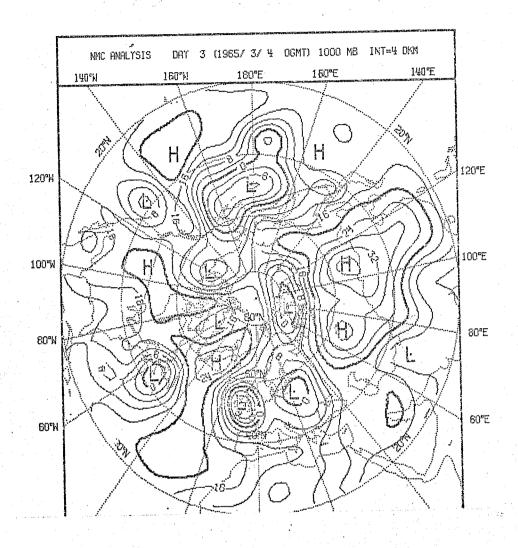
FIGURE 13

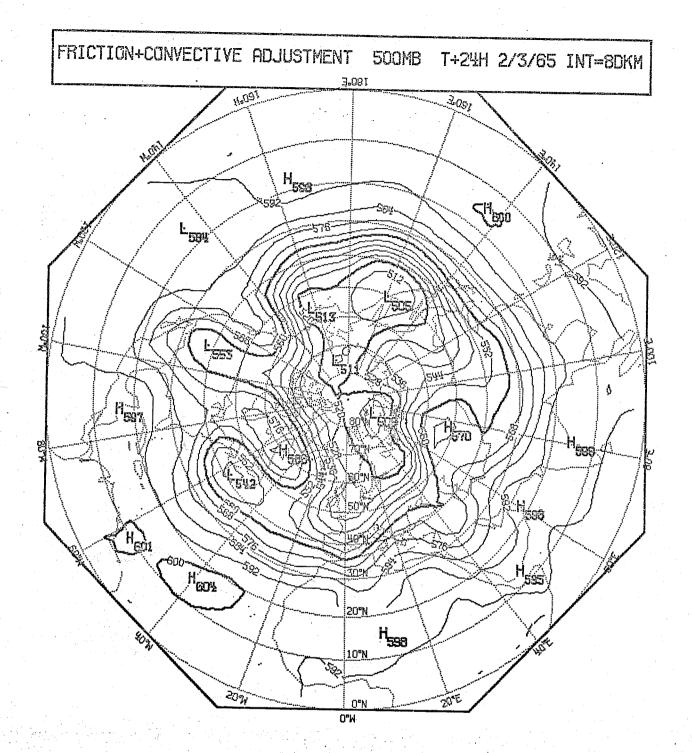


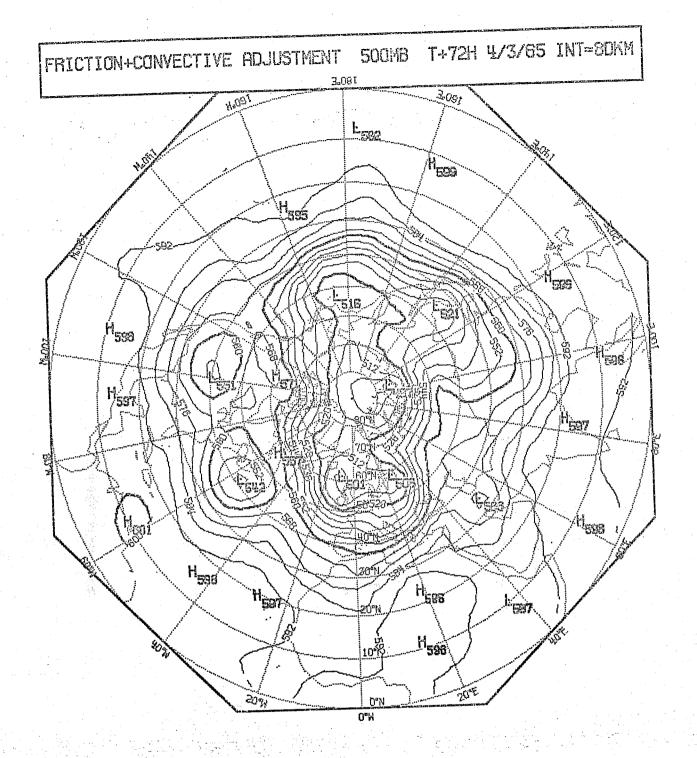


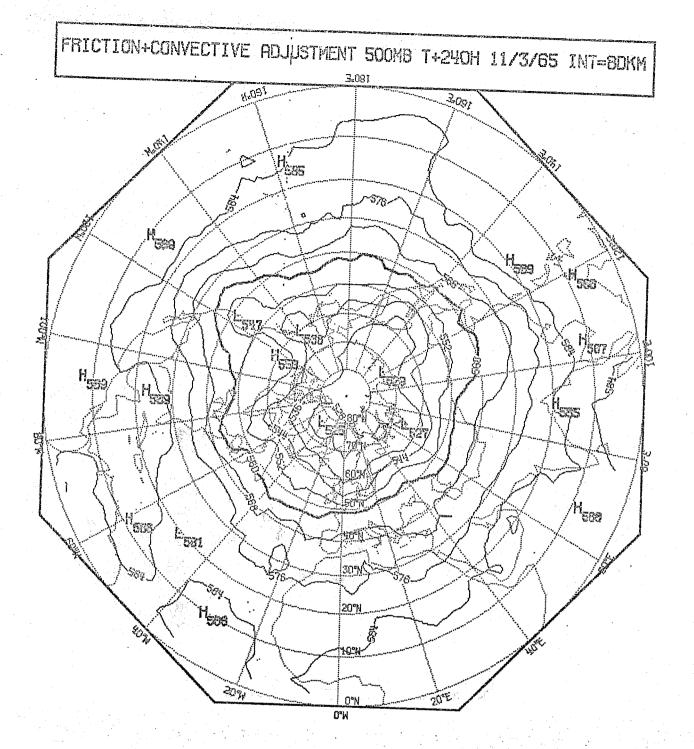




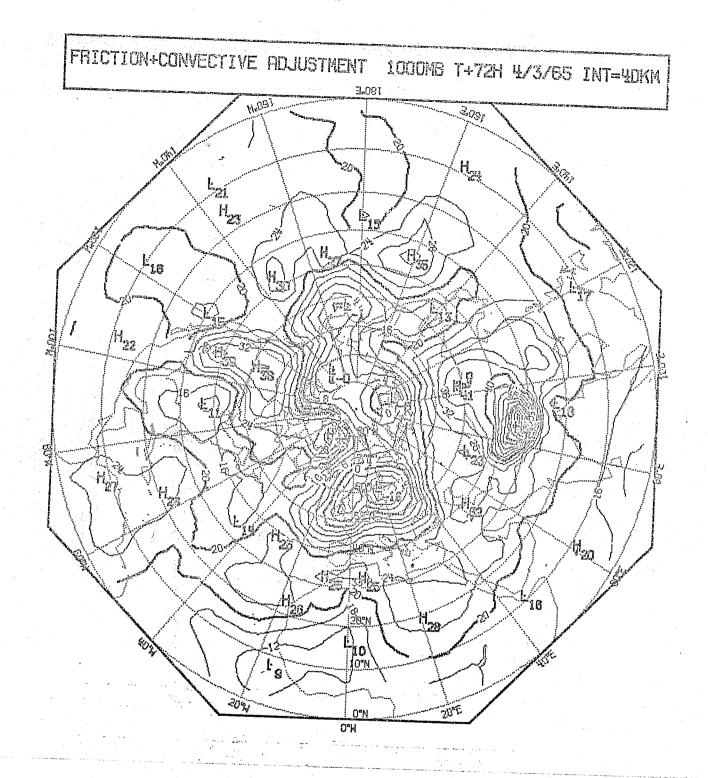


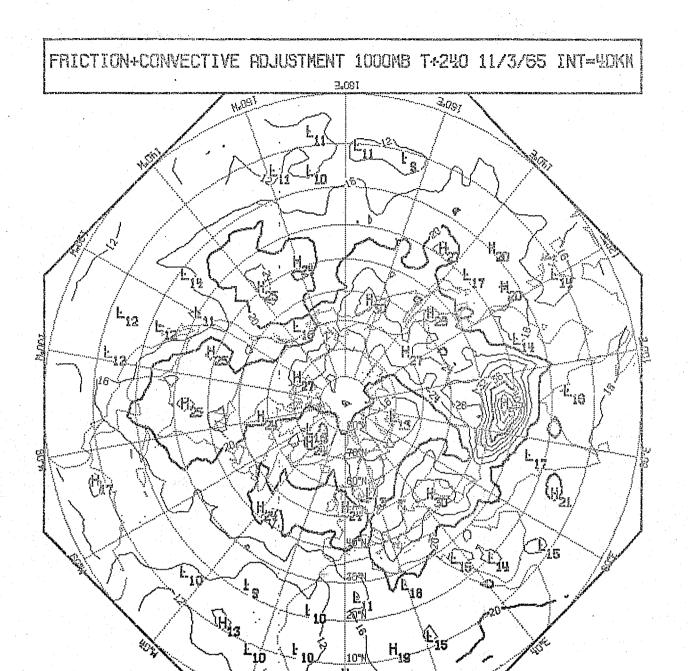






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